

Code-Carrier Divergence Monitoring for the GPS Local Area Augmentation System

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Abstract

Code-carrier smoothing is a commonly used method in Differential GPS (DGPS) systems to mitigate the effects of receiver noise and multipath. The FAA's Local Area Augmentation System (LAAS) uses this technique to help provide the navigation performance needed for aircraft precision approach and landing. However, unless the reference and user smoothing filter implementations are carefully matched, divergence between the code and carrier ranging measurements will cause differential ranging errors.

The FAA's LAAS Ground Facility (LGF) reference station will implement a prescribed first-order Linear Time Invariant (LTI) filter. Yet flexibility must be provided to avionics manufacturers in their airborne filter implementations. While the LGF LTI filter is one possible means for airborne use, its relatively slow transient response (acceptable for a ground based receiver) is not ideal at the aircraft because of frequent filter resets following losses of low elevation satellite signals (caused by aircraft attitude motion). However, in the presence of a code-carrier divergence (CCD) anomaly at the GPS satellite, large divergence rates are theoretically possible, and therefore protection must be provided by the LGF through direct monitoring for such events. In response, this paper addresses the impact of the CCD threat to LAAS differential ranging error and defines an LGF monitor to ensure navigation integrity.

Differential ranging errors resulting from unmatched filter designs and different ground/air filter start times are analyzed in detail, and the requirements for the LGF CCD monitor are derived. A CCD integrity monitor algorithm is then developed to directly estimate and detect anomalous divergence rates. The monitor algorithm is implemented and successfully tested using archived field data from the LAAS Test Prototype (LTP) at the William J. Hughes FAA Technical Center. Finally, the paper provides recommendations for initial monitor implementation and future work.

INTRODUCTION

The Federal Aviation Administration (FAA) and the aviation industry are currently developing the Local Area Augmentation System (LAAS). This is a differential GPS system that augments GPS navigation in two ways. First, the LAAS Ground Facility (LGF) provides differential corrections to the user (aircraft) that augment user accuracy. Second, the LGF monitors the ranging sources to protect against faults that could result in navigation errors, thereby ensuring the integrity of navigation for LAAS users.

In order to ensure system integrity, for each class of ranging

source (i.e., satellite) fault there is a monitor at the LGF that computes a test metric designed to indicate the presence of that particular type of fault. The purpose of these monitors is to limit the LAAS user's integrity risk due to a fault occurring at the ranging source.

LAAS ranging source faults are divided into two categories [2]: an additive bias to the differential correction error with no change in variance of the error, and an increase in variance of the differential correction error with no change in bias. Faults that result in the first kind of error are Code-Carrier Divergence (CCD), signal deformation, erroneous ephemeris, and excessive acceleration. A satellite low power fault results in an increase in the ranging error variance leading to a fault of the second type.

In this paper we address the code-carrier divergence fault. Divergence can be caused by ionospheric activity (nominal or storm) or a fault occurring at the ranging source (satellite). The latter is the primary CCD threat, as large divergence rates are theoretically possible. This motivates the need for CCD monitoring to be provided by the LGF. During ionospheric activity, the CCD monitor can also provide benefit by helping to detect moving storm fronts. However, the LGF has other monitors designated to detect ionospheric anomalies.

Like other DGPS architectures, the LGF (reference station) and aircraft (user) employ code-carrier smoothing to mitigate the effects of receiver noise and multipath. If the aircraft and ground implement identical filter designs and have the same start times, then they experience the same transient response to divergence, and differential-ranging errors will not exist. However, if they have unmatched filter designs or different start times, then divergence between code and carrier ranging measurements will cause differential ranging errors.

The LGF reference station will implement an FAA prescribed first-order Linear Time Invariant (LTI) filter. One approach to mitigate the CCD threat is to have the ground and aircraft implement identical smoothing filters. The cost, however, is a loss of design flexibility for avionics manufactures, which is not desirable. Even though the LGF LTI filter is one possible means for airborne use, its relatively slow transient response (acceptable for a ground based receiver) is not ideal at the aircraft because the aircraft filter is likely to experience frequent filter resets following the loss of signals from low elevation satellites (caused by aircraft attitude motion). In this

paper we will consider various possible aircraft filter implementations and show that a first order Linear Time Varying (LTV) filter is a good choice for airborne implementation.

The CCD monitor algorithm described in this paper has two functions: divergence rate estimation and detection. Monitor thresholds are established with the aid of LAAS Test Prototype (LTP) data. For the integrity analysis we present a new direct approach for computing the overall probability of Loss Of Integrity $P(LOI)$ and describe a preliminary analysis for time to alert.

The paper is divided into five sections. Section I describes the relevant requirements for CCD fault detection. Section II discusses the different types of possible smoothing filters that can be implemented at the aircraft. Section III defines the CCD monitor and provides experimental validation using archived field data from the LAAS Test Prototype (LTP) at the William J. Hughes FAA Technical Center and LGF test data provided by Honeywell. Section IV presents the general integrity analysis for a space segment failure with application to CCD monitoring. Finally, Section V gives the conclusion.

I. REQUIREMENTS

This section briefly summarizes the relevant LGF specifications and requirements on ranging source integrity, ground monitor continuity, and the prescribed LGF smoothing filter algorithm. The relevant Minimal Operational Performance Standards (MOPS) avionics requirements are also discussed. Further details on requirements can be found in [4, 5].

A. LGF Category I Integrity Requirement: The probability that the LGF transmits Misleading Information (MI) for 3 seconds or longer due to a Ranging Source (RS) failure shall not exceed 1.5×10^{-7} during any 150 sec approach interval.

- MI is defined as broadcast data that results in the lateral or vertical position error exceeding protection levels for any user w/in 60 nmi of the LGF.
- The CCD failure rate is defined to be 10^{-4} /hr for satellite (SV) acquisition, and the prior probability of CCD failure after acquisition is given as 4.2×10^{-6} /SV/approach.

B. LGF Reference Receiver (RR) and Ground Monitor Continuity Requirement: It is required that the probability of any valid ranging source is made unavailable due to a false alarm shall not exceed 2.3×10^{-6} per 15 sec interval.

C. Requirement Allocations: For the purposes of this work, we assume that the relevant allocations (from items A and B above) for the CCD monitor are:

- The probability of MI given a CCD fault is 10^{-4} per 150 sec approach interval, and the probability of fault free alarm is 10^{-7} per 15 sec interval.

D. LGF Smoothing Filter Algorithm: The smoothing filter is defined in the LGF specification as follows. In steady state, each pseudorange measurement from each RS shall be smoothed using the filter:

$$PR_s = \left(\frac{1}{N} \right) PR_r(k) + \left(\frac{N-1}{N} \right) [PR_s(k-1) + \phi(k) + \phi(k-1)]$$

$$N = \frac{S}{T}$$

where, PR_r = raw pseudorange,

PR_s = the smoothed pseudorange,

S = time filter constant, equal to 100 seconds,

T = filter sample interval, nominally equal to 0.5 seconds,

ϕ = accumulated phase measurement,

k = current measurement, and

$k-1$ = previous measurement.

The LGF is also required to generate and broadcast to LAAS users σ_{pr_gnd} , the standard deviation of the error on the broadcast smoothed pseudorange correction. The variable σ_{pr_gnd} can be expressed as the root sum square (RSS) of the standard deviation of ground ranging error due to all sources except filter transient to nominal ionospheric divergence ($\sigma_{pr_gnd,nom}$) and the standard deviation that accounts for ground transient filter responses to nominal ionospheric divergence (σ_{div_gnd}). The nominal ionospheric CCD rate is given in LGF Specification as normally distributed with zero mean and standard deviation of 0.018 m/s.

E. MOPS Avionics Requirements: The airborne system is also required to do carrier smoothing of the pseudorange measurements, but a specific filter is not defined. (This differs from the LGF Specification, which prescribes the ground system filter implementation). The MOPS provides significant flexibility to avionics manufacturers by specifying that the airborne filter need only match the ground filter with the following performance requirement:

“In response to a code-carrier divergence rate of up to 0.018 m/s, the smoothing filter output shall achieve an error less than 0.25 m within 200 sec after initialization relative to the steady-state response of [filters specified in LGF Spec].”

Like the ground system, the aircraft also has a σ_{pr_air} which can be expressed as the RSS of $\sigma_{pr_air,nom}$ (standard deviation of air ranging error due to all sources except filter transient to nominal ionospheric divergence) and σ_{div_air} (standard deviation of air ranging error due to filter transient to nominal ionospheric divergence). The MOPS states that the aircraft must account for its filter transient response to nominal divergence (i.e., after filter startup or reset) by inflating the standard deviation used for ranging measurements in the derivation of its protection levels. This effect is captured by σ_{div_air} .

The MOPS further requires that the steady state value of σ_{div_air} shall not exceed 0.15 m, and it states that “steady state operation is defined to be following 360 seconds of continuous operation of the smoothing filter.”

The relevant MOPS requirements can be summarized in graphical form as shown in Figure 1. The aircraft is permitted to have transient response to divergence of 0.018 m/s anywhere in the unshaded region. The response of the LGF filter (one acceptable choice at the aircraft) is shown in the Figure 1.

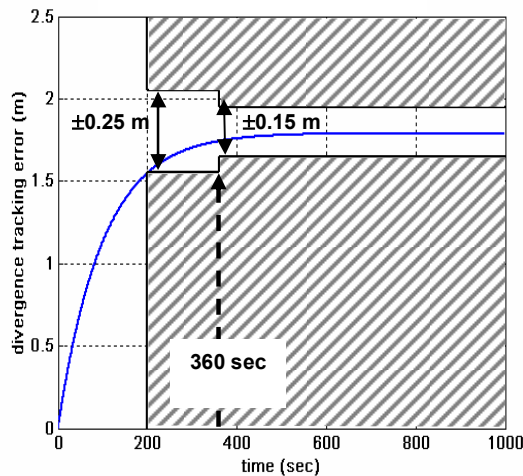


Figure 1. MOPS-Compliant Transient Response

II. GENERAL AIRCRAFT FILTERS

From Figure 1, it is clear if the order and time-invariance of the avionics filter are unconstrained, there exists flexibility in MOPS-compliant transient response to divergence inputs. Unfortunately, any significant variation from the LGF filter response is problematic (even if permitted by the MOPS) because it will result in a differential ranging error. The aircraft must account for any such deviation in σ_{div_air} , which increases σ_{pr_air} as it is defined earlier—i.e. RSS ($\sigma_{pr_air_nom}$, σ_{div_air}), with $\sigma_{pr_air_nom}$ being the nominal value of σ_{pr_air} prior to accounting for ionosphere divergence effects at the aircraft. But increasing σ_{pr_air} can reduce system availability. Thus, in order to maximize system availability, it is reasonable to assume that the goal of the avionics manufacturer is to keep σ_{pr_air} small at any given time. For divergence inputs, this means that minimizing overshoot and steady-state error will lower σ_{pr_air} . For nominal code noise inputs, a time varying filter implementation (at filter start-up) can provide quicker noise response, which will lower $\sigma_{pr_air_nom}$ and therefore σ_{pr_air} . So these types of filters need to be considered. In this section, we show practical first and second order avionics filter designs that are potentially useful.

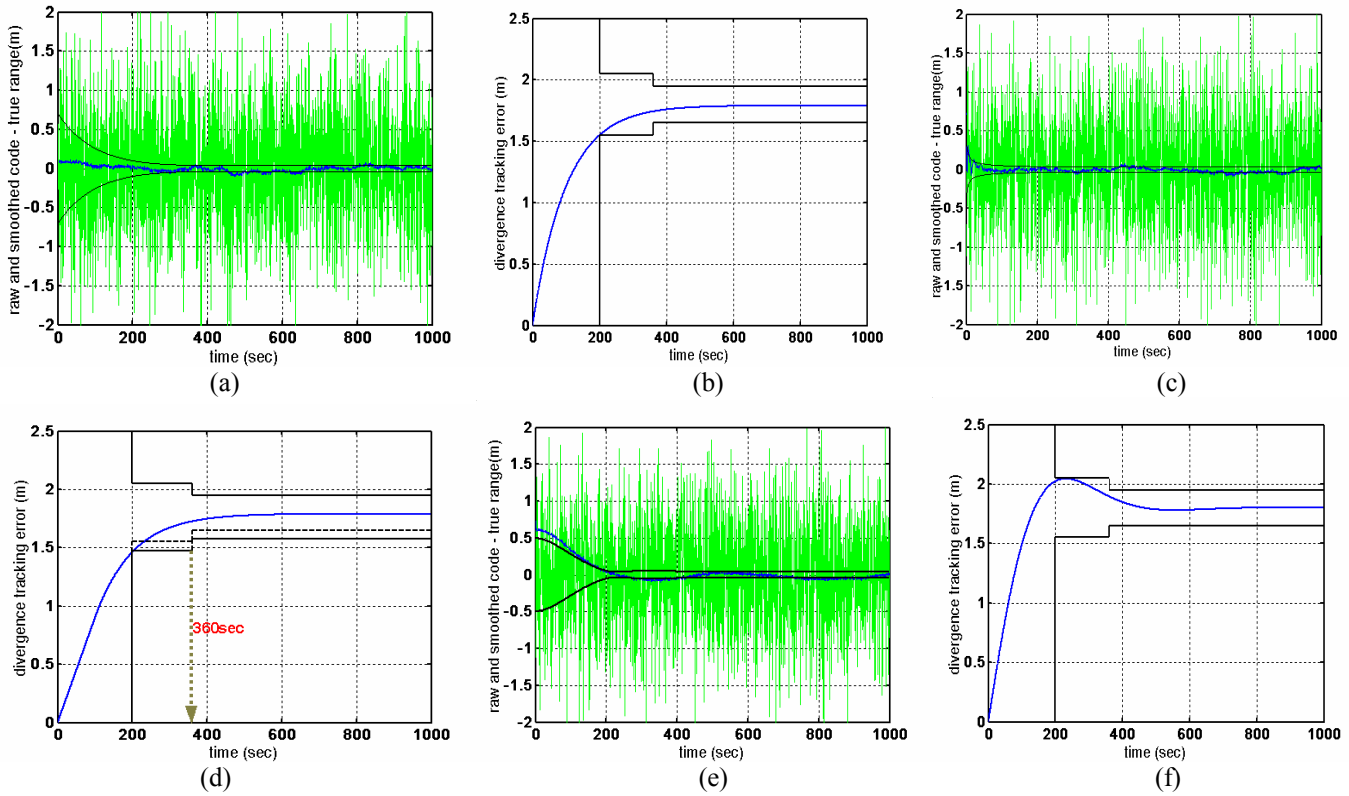
Figures 2(a) through Figure 2(d) show the transient responses of 1st order LTI and LTV MOPS-compliant aircraft filter implementations. Figure 2(a) shows the sample response to a white code noise input ($\sigma = 0.5$ m) for the LGF LTI filter

along with the theoretical 1-sigma error envelope. Figure 2(b) shows the response of the same filter to a nominal divergence input. The slow transient response of this filter to noise will affect the aircraft more seriously than the ground because of frequent filter resets at the aircraft.

Using a time varying gain in the first order filter at the aircraft can significantly speed up the transient response to noise. Figure 2(c) shows a sample noise response using a filter with time constant $\tau = t$ for $t < 100$ sec and $\tau = 100$ sec for $t \geq 100$ sec. The rapid convergence of the error is readily apparent as compared to the LTI case. Furthermore, the filter is simple to implement as it deviates from the LGF LTI filter only in that the gains are changing in the first 100 sec. The response of the LTV filter to a nominal divergence input of 0.018 m/s, shown in Figure 2(d), suggests that the MOPS divergence response requirements are not met (dashed lower tolerable limit line). However, when the MOPS definition of steady state (360 sec) is used to define the lower transient response boundary (solid line), the time response becomes acceptable. From the point of view of LGF CCD monitor design the LTV filter must therefore be treated as a realistic candidate for airborne implementation.

The use of second order filters at the aircraft can also be MOPS compliant. Figure 2(e) and Figure 2(f) show the noise and divergence responses of a 2nd order LTI filter, which was designed to simultaneously maximize overshoot (subject to the MOPS divergence response requirements in Figure 1) and yet produce negligible steady state error relative to the LGF implementation. As with the first order case, second order LTV realizations are also possible to speed up noise response. While there is no clear time response benefit in the use of 2nd order filters in the avionics relative to 1st order LTV filter, the noise output of the 2nd order filter does have less high frequency content. This may be beneficial to a navigation avionics manufacturer who desires to produce smoother position inputs to the autopilot. Although, 2nd order filters do not help in lowering σ_{pr_air} , the potential benefit of lowering the high-frequency content of the output means that the potential for airborne implementation of such a filter cannot be ignored. The overshoot exhibited in Figure 2(f) is not a necessary feature of a second order implementation, but it does define the MOPS compliant upper limit on the transient response for such filters. The overshoot in this 2nd order response would cause a temporary increase in σ_{pr_air} , (relative to a 1st order implementation) which may be acceptable to an avionics manufacturer. In any case, however, it is reasonable to assume that the aircraft would not intentionally implement a filter with a steady state error different from the reference LGF filter, as it would cause an unnecessary increase in σ_{pr_air} for the entire satellite pass therefore cause a potential availability penalty.

As discussed earlier, the MOPS gives enough room to implement a second order LTI/LTV filter. However, for the



rest of the analysis in this paper we assume that the aircraft uses a first order LTV filter for smoothing. The implication on results if a second order filter is used can be considered for future work. In the next section, we present the algorithm development of the CCD monitor.

III. CCD Monitor Algorithm

The LGF divergence monitor consists of two components: a *divergence rate estimator* and a *detection test*. The input to the divergence rate estimator, z , is the raw code minus carrier measurement. The divergence rate estimator differentiates the input z and filters the result using two first order LTI filters in series (second order filter with real poles) to reduce the code noise contribution to the estimation error. The continuous-time realization of the estimator algorithm (Laplace Transform) is,

$$\hat{D}(s) = \frac{s}{(\tau_{d1}s + 1)(\tau_{d2}s + 1)} Z(s) \quad (1)$$

The estimator output (\hat{d} in time domain) is the filtered divergence estimate. The discrete time equations for the rate estimator are given below with $d_2(k)$ being the final rate estimate:

$$\left. \begin{aligned} d_1(k) &= \frac{\tau_{d1} - T}{\tau_{d1}} d_1(k-1) + \frac{1}{\tau_{d1}} [z(k) - z(k-1)] \\ d_2(k) &= \frac{\tau_{d2} - T}{\tau_{d2}} d_2(k-1) + \frac{T}{\tau_{d2}} d_1(k-1) \end{aligned} \right\} \quad (2)$$

The filter time constants τ_{d1} and τ_{d2} are algorithm parameters, whose values are nominally set at $\tau_{d1} = \tau_{d2} = 30$ sec. T is the sample time. The justification for these particular time constant values will be provided shortly.

The use of two 1st order filters in series reduces estimate error caused by differentiating high-frequency code measurement noise when compared to a single 1st order filter. This effect is clearly demonstrated in Figure 3 and Figure 4, which show filter outputs to simulated white noise with standard deviation of 0.5 m along with the theoretically derived standard deviation envelopes. It is clear from these two figures that the performance of the 2nd order implementation with a 30 sec time constant is superior (lower output noise) to a 1st order implementation, even when the latter has a much longer filter time constant (200 sec). The ability to use a shorter time constant will result in quicker detection of CCD failures.

Following the divergence estimator is the detection function, which is a simple threshold test: If $d_2 > T_{ccd}$ then alarm, else no alarm. The detection threshold defined as:

$$T_{ccd} = k_{ffd,mon} \sigma_d \quad (3)$$

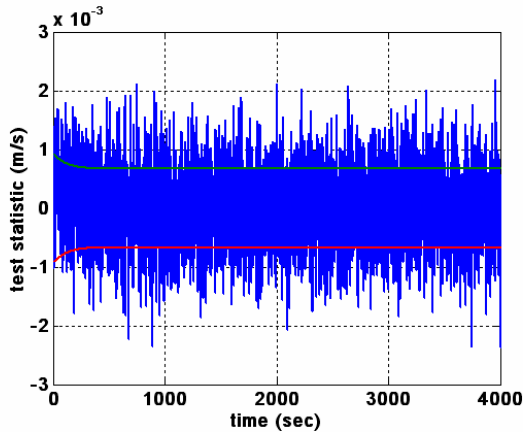


Figure 3. 1st Order Filter (\hat{d} response) $\tau_{d1} = 200\text{sec}$.

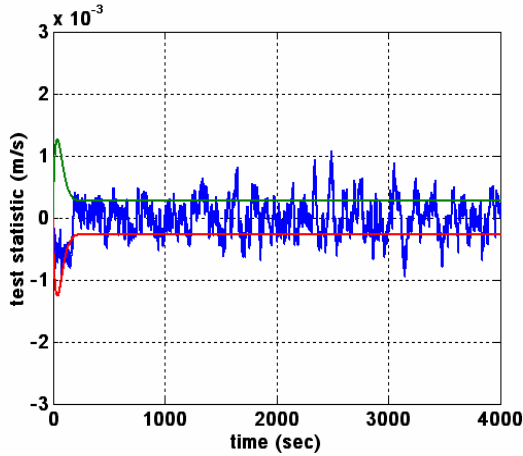


Figure 4. Two 1st Order Filters in series ($\tau_{d1} = \tau_{d2} = 30\text{sec}$)

In equation (3), σ_d is the fault-free standard deviation of the test statistic d_2 and $k_{ffd,mon}$ is a constant chosen to ensure that the probability of fault-free alarm meets the allocated continuity requirement for the monitor [3]. Both σ_d and $k_{ffd,mon}$ are algorithm parameters, whose values are nominally set at $\sigma_d = 0.00399$ m/s and $k_{ffd,mon} = 5.83$. The justification for σ_d value is given in the next sub-sections.

A. Fault-Free Distribution: As stated above, in order to set the monitor threshold, we need the fault-free standard deviation (σ_d) of test statistic d_2 . This fault-free distribution will be affected by both filtered code noise (σ_{dn}) and nominal ionospheric divergence (σ_{di}) which are assumed to have distributions $N(0, \sigma_{dn})$ and $N(0, \sigma_{di})$ respectively, such that:

$$\sigma_d = \sqrt{\sigma_{dn}(EL, \tau_{d1}, \tau_{d2})^2 + \sigma_{di}^2} \quad (4)$$

where, EL is the satellite elevation. We will first consider the nominal ionospheric contribution.

1) *Nominal Ionospheric Divergence Contribution:* The nominal value for ionospheric divergence used in the LGF specification and the MOPS ($\sigma_{di} = 0.018$ m/s) was selected to ensure that the protection levels computed at the aircraft would be conservatively large. However, assuming such a large value for σ_{di} for the CCD monitor will cause unreasonably loose thresholds. In fact, prior research [10] suggests that a more realistic nominal value is $\sigma_{di} \approx 0.003$ m/s. Archived dual frequency carrier phase data from the LTP was used to substantiate this result, as is discussed below.

Ashtech receiver L1/L2 carrier phase data was used to compute instantaneous ionospheric divergence rates over 1 sec measurement intervals with a 5 deg elevation mask. The instantaneous rates were then averaged in 100 sec windows to reduce the differentiated carrier phase noise contribution to the rate estimates. The analysis was carried out for 17 satellites over 7 months with data taken from one day in each of these months (see Table 1). Figure 5 shows an example divergence trace for a single satellite on a single pass.

Table 1. Nominal Ionospheric Divergence Data Archive

SV #	Date (Year 2004)
1, 3, 4, 5, 6, 7, 8, 9, 10, 11, 13, 14, 15, 16, 17, 18, 20	Feb 11, Mar 11, Apr 16, Jun 15, Jul 15, Aug 31, and Oct 05

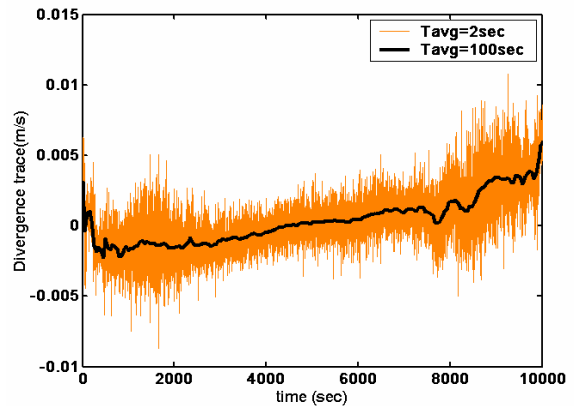


Figure 5. Example Divergence trace: SV#3 Feb 11 '04.

Figure 6 shows the cumulative distribution function (CDF) of all of the empirical divergence rates from the entire data set. The dashed line in the plot is the CDF of a gaussian distribution with the same standard deviation as the data ($\sigma_{di} = 0.0014$ m/s). It is clear that the ionospheric divergence rate distribution is not gaussian in the distribution tails. However, Figure 7 shows that when the original standard deviation is inflated by a factor of 2.85, the new gaussian distribution does bound the empirical CDF.

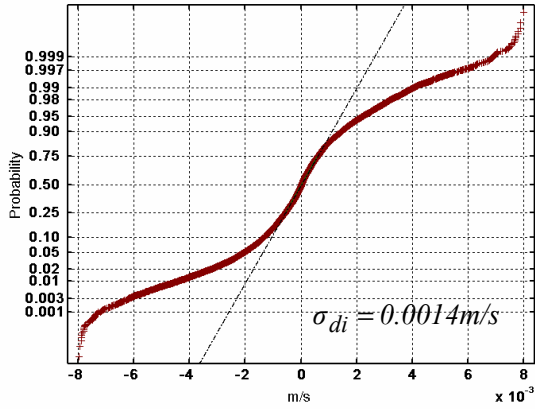


Figure 6. CDF Plot for Combined LTP Divergence Data

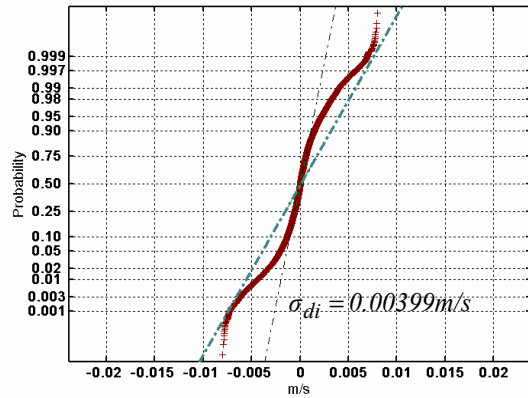


Figure 7. Gaussian Overbound of LTP Ionospheric Divergence Rate Data.

Thus, the value to be used is $\sigma_{di} = 2.85 \times 0.0014 = 0.00399$ m/s.

2. Filtered Code Noise Contribution: As indicated in equation (4), the contribution of filtered code noise (σ_{dn}) depends on satellite elevation (EL) and the filter time constants τ_{d1} and τ_{d2} . In principle, τ_{d1} and τ_{d2} can always be chosen large enough to make $\sigma_d \approx \sigma_{di}$ (although care must be taken to ensure that the time to detect does not become too large). For simplicity in the monitor implementation and analysis, we define $\tau_d = \tau_{d1} = \tau_{d2}$.

Archived field test data provided by Honeywell International (LGF contractor) was used to determine σ_{dn} as a function of EL for different values of τ_d . Data from the Honeywell LGF receivers equipped with Multipath Limiting Antennas (MLA) and High Zenith Antennas (HZA) was used as the source for code and carrier data as input into the divergence rate estimator in equation (2). To isolate the code noise contribution, carrier phase data from a nearby dual frequency Novatel OEM4 receiver was used to remove nominal ionospheric divergence from the data prior to processing. Table 2 provides a summary of the empirical results for σ_{dn} in m/sec.

It is clear from the results that the resulting σ_{dn} is an order of magnitude smaller than σ_{di} . The contribution of σ_{dn} to σ_d is negligible even for $\tau_d = 20$ sec relative to the effect of σ_{di} . Figure 8 shows the standard deviation of filtered code noise against time constant for a single satellite.

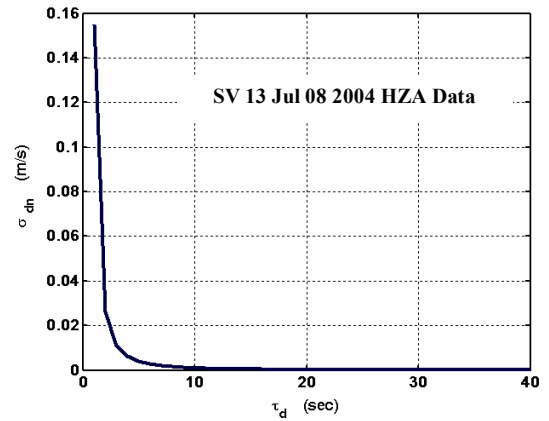


Figure 8. HZA Divergence Estimate error vs. Time constant

For this initial monitor analysis a nominal value for $\tau_d = \tau_{d1} = \tau_{d2} = 30$ sec is selected.

Table 2. σ_{dn} (m/s) vs. Elevation and Time Constant

τ_d \ EL	EL					
	5-15 (MLA)	15-30 (MLA)	30-40 (HZA)	40-50 (HZA)	50-60 (HZA)	>60 (HZA)
20sec	0.00038	0.00035	0.0003	0.0002	0.0001	0.0001
40sec	0.00010	0.000096	0.000074	0.00004	0.000034	0.000032
60sec	0.000072	0.000045	0.00006	0.00002	0.000015	0.000015

Table 3. Monitor Modifiable Parameters

Monitor Parameter	Definition	Value
T	Sample Time	0.5sec
$k_{ffd,mon}$	Constant	5.83
τ_d	Time Constant	30sec
σ_d	std. dev. of d_2 (fault free test metric)	0.00399 m/s

Having set the nominal CCD monitor parameters, which are summarized in Table 3, we next proceed to the integrity analysis.

IV. General Integrity Analysis for Space Segment Failure

The following definitions will be used below in the derivation of the probability of Loss of Integrity (LOI) due to a general space segment failure event on satellite k :

- e_v Vertical component of differential position error due to all sources
- v_i Differential ranging error for satellite i due to all fault-free error sources
- σ_i Standard deviation of v_i
- b_k Differential ranging error on satellite k due to satellite failure only
- n Number of satellites in view
- S Weighted pseudoinverse used at aircraft for position fix (which is a function of satellite geometry)
- $S_{v,i}$ The element of matrix S projecting the ranging measurement from satellite i into the vertical direction
- q_k Test statistic for the fault monitor for satellite k due to all sources
- r_k Test statistic for the fault monitor for satellite k due to the satellite failure only
- η_k Test statistic for the fault monitor for satellite k due to all fault-free error sources
- $\sigma_{\eta,k}$ Standard deviation of the monitor test statistic in the fault-free case
- $\sigma_{r,k}$ Standard deviation of the monitor test statistic for faulted case
- $k_{ffd,mon}$ The multiplier on $\sigma_{\eta,k}$ used to define monitor threshold with desired fault free detection probability
- Q Cumulative distribution function for the standard normal distribution

$\sum_{i=1}^n S_{v,i} v_i$ Vertical component of differential position error for all fault-free error sources

σ_v Standard deviation of $\sum_{i=1}^n S_{v,i} v_i$

k_{ffd} 5.81, the multiplier on σ_r used to compute LAAS VPL_{H0}

In this analysis, the probability of Loss of Integrity defined as:

$$P(LOI|fault_k) \equiv P(|e_v| > VPL|fault_k) P(|q_k| < k_{ffd,mon} \sigma_{r,k} | fault_k) \quad (5)$$

Consider the two terms on the right-hand side of equation (5) separately. Given a failure on satellite k , with a resulting b_k not close to zero, the first term on the right hand side can be simplified as a one-sided probability (Figure 9) given by:

$$\begin{aligned} P(|e_v| > VPL | fault_k) &\approx P(|S_{v,k} b_k| + \sum_{i=1}^n S_{v,i} v_i > k_{ffd} \sigma_v) \\ &= 1 - Q\left\{k_{ffd} \sigma_v - |S_{v,k} b_k|\right\} \\ &= 1 - Q\left\{k_{ffd} - \frac{|S_{v,k} b_k|}{\sigma_v}\right\} \\ &= 1 - Q\left\{k_{ffd} - \frac{\sigma_{v,k} |b_k|}{\sigma_v \sigma_k}\right\}. \end{aligned}$$

The worst-case probability occurs when $\frac{\sigma_{v,k}}{\sigma_v} \rightarrow 1$. Therefore, it is convenient to conservatively use the following satellite-

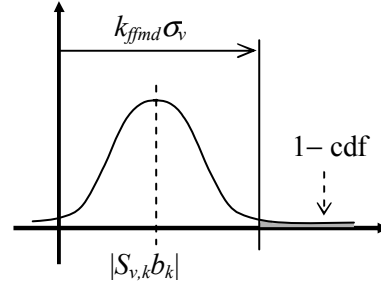


Figure 9. LGF Integrity Risk given fault on Ranging Source k

geometry-free expression,

$$\begin{aligned} P(|e_v| > VPL | fault) &= 1 - Q\left\{k_{ffd} - \frac{|b_k|}{\sigma_k}\right\} \\ &= Q\left\{\frac{|b_k|}{\sigma_k} - k_{ffd}\right\} \quad (6) \end{aligned}$$

Using the same method the second term on the right-hand side of equation (5) can be simplified as follows:

$$\begin{aligned} P(|q_k| < k_{ffd,mon} \sigma_r | fault_k) &\approx P(|r_k| + \eta_k < k_{ffd,mon} \sigma_r | fault_k) \\ &= Q\left\{k_{ffd,mon} - \frac{|r_k|}{\sigma_{r,k}}\right\} \quad (7) \end{aligned}$$

Substituting the results (6) and (7) in to equation (5) yields,

$$P(LOI | fault_k) = Q\left\{\frac{|b_k|}{\sigma_k} - k_{ffnd}\right\} Q\left\{k_{ffd,mon} - \frac{|r_k|}{\sigma_{r,k}}\right\} \quad (8)$$

In general, a space segment failure event on satellite k will cause different transient responses in the differential position error b_k and the monitor test statistic r_k . The loss of integrity probability in equation (8) will be a function of both of these failure response functions, the failure magnitude, the elapsed time since failure onset, and the ground and airborne receiver tracking start times (which influence σ_k and $\sigma_{r,k}$). For every type of satellite failure it is necessary to find the conditions that maximize the LOI probability. It is also necessary to determine whether the LOI probability exceeds the integrity risk allocation for a failure mode for duration greater than the maximum permitted time-to-alert. For LGF integrity monitors, the required time-to-alert is 3 sec.

A. Application to CCD Monitoring: In the section the goal is to obtain the terms in equation (8), when a failure occurring at the ranging source causes a CCD fault. In order to do so we define the following:

t_{0g} LGF filter start time

t_{0a} Airborne filter start time

$d_{nom} = 0.018$ m/s, the nominal ionospheric divergence rate

e_g , Ranging error due to d_{nom} at LGF relative to LGF steady state

e_a , Ranging error due to d_{nom} at aircraft relative to LGF steady state

Given a divergence failure with a CCD rate d and time of onset $t = 0$, a differential ranging error exists only when both filters are tracking ($t > t_{0a}$ and $t > t_{0g}$) and after onset of the failure ($t > 0$). Therefore the differential ranging error, b_k in the general analysis in the preceding section, can be expressed as

$$b_k = \frac{d}{d_{nom}} \left| e_g(t - \max[t_{0g}, 0]) - e_a(t - \max[t_{0a}, 0]) \right| \quad (9)$$

Note that t_{0g} and t_{0a} can be negative, signifying possible filter start times prior to failure onset.

The standard deviation of the ranging error used at the aircraft, σ_k in the preceding section, is

$$\sigma_k = \sqrt{\sigma_{pr_gnd}^2 + \sigma_{div_gnd}^2 + \sigma_{pr_air}^2 + \sigma_{div_air}^2} \quad (10)$$

As $d_{nom} = 0.018$ m/s is prescribed by the LAAS MOPS and LGF Specification to be used as the standard deviation of

nominal ionospheric divergence rate, then $e_g = \sigma_{div_gnd}$ and $e_a = \sigma_{div_air}$. Therefore,

$$\sigma_k = \sqrt{\sigma_{pr_gnd}^2 + \sigma_{pr_gnd}^2 + e_g(t - t_{0g})^2 + e_a(t - t_{0a})^2} \quad (11)$$

and combining with (9),

$$\frac{b_k}{\sigma_k} = \frac{d \left| e_g(t - \max[t_{0g}, 0]) - e_a(t - \max[t_{0a}, 0]) \right|}{d_{nom} \sqrt{\sigma_{pr_gnd}^2 + \sigma_{pr_air}^2 + e_g(t - t_{0g})^2 + e_a(t - t_{0a})^2}} \quad (12)$$

Equation (12) may be substituted into the first term on right-hand side of equation (8) for the CCD case. Ideally, the LGF monitor performance should be independent of $\sigma_{pr_gnd,nom}$ and $\sigma_{pr_air,nom}$, which may vary over time and location. In the most conservative analysis, it is assumed that $\sigma_{pr_gnd,nom} = \sigma_{pr_air,nom} = 0$.

As discussed earlier, the LGF filter is a first order digital LTI filter with a 100 sec time constant (defined in the LGF Specification). In the current LGF prototype implementation, there is no correction broadcast for satellites during the first 200 sec of filtering (i.e., only $t - t_{0g} > 200$ sec needs to be considered in the integrity analysis). For the aircraft, as discussed in section II, it is assumed a first order digital LTV filter is implemented. This filter differs from the LGF filter only during the first 100 sec of operation, when the effective filter time constant increases uniformly in time (up to the 100 sec limit). It is also assumed that the aircraft will use filtered measurements immediately (i.e., $t - t_{0a} > 0$ needs to be considered in the integrity analysis). The two digital filter responses to nominal divergence, e_g and e_a , are plotted as function of time in Figure 10.

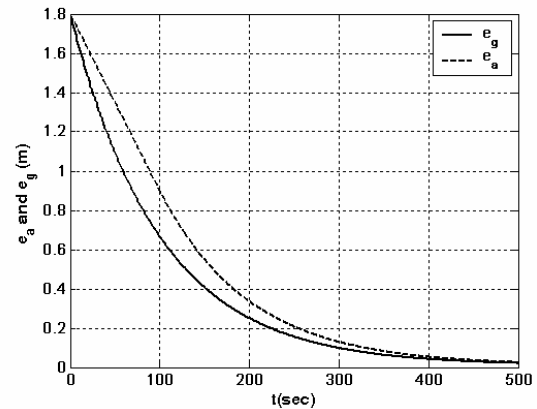


Figure 10. Differential ranging error relative to LGF steady state

For the CCD monitor, the divergence rate estimate d_2 is the test statistic r_k in the general analysis in the preceding section. Therefore,

$$\frac{r_k}{\sigma_{r,k}} = \begin{cases} \frac{|d_2(t)|}{0.00399 \text{ m/s}} & t > t_{0g} \\ 0 & t < t_{0g} \end{cases} \quad (13)$$

This may be substituted into the second term on the right-hand side of equation (8) for the CCD case.

The example noise-free time response of d_2 to a unit ramp divergence input, $d = 1$, is shown in Figure 11 for the digitally implemented estimator. Because the divergence estimator is a linear filter, the amplitude of the time response for other divergence inputs will simply scale linearly with d .

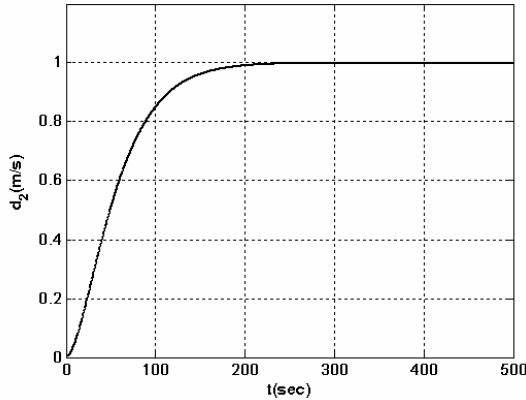


Figure 11. Noise free Divergence Estimate for $d=1\text{m/s}$.

B. Integrity Analysis Results: The divergence onset of magnitude d is defined to occur at time $t = 0$. For the purpose of interpreting the integrity analysis results, we will refer to the first term on the right-hand side of equation (8) as $P_{ev|fault}(t, t_{0a}, t_{0g}, d)$ and the second term as $P_{md|fault}(t - \max[t_{0g}, 0], d)$.

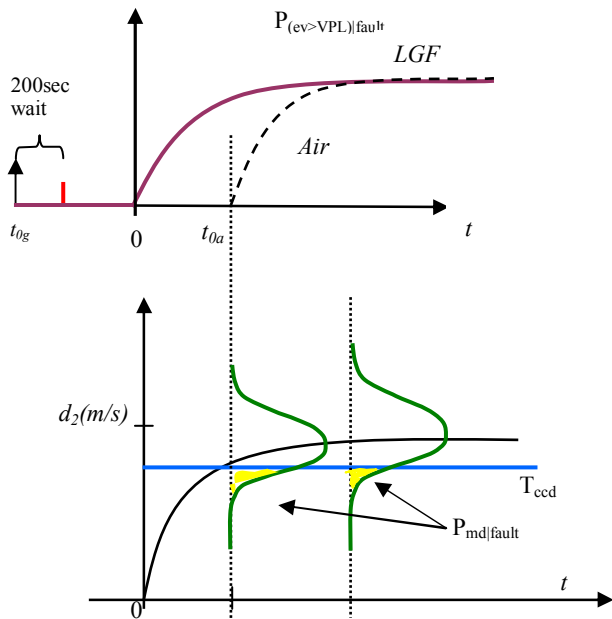


Figure 12. Illustration for $P_{(ev|fault)}$ and $P_{md|fault}$

If the ground filter starts close in time to the aircraft filter, then the monitor response for the first 200 seconds is not utilized for computation of $P_{md|fault}$ because the LGF does not broadcast corrections during this time. Hence $P_{md|fault} = 0$ during the first 200 sec. Therefore, as illustrated in Figure 12 the worst-case $P(LOI|fault_k)$ occurs when the ground filter has started well before the aircraft filter: theoretically speaking, when $t_{0g} = -\infty$ and aircraft filter has just started. This result is further illustrated in Figure 13, which shows contours of the highest values of $P(LOI|fault_k)$ for any value of t at specified values of t_{0g} and t_{0a} . The highest values of $P(LOI|fault_k)$ occur where in t_{0g} is at its lowest value (-500 sec in the figure).

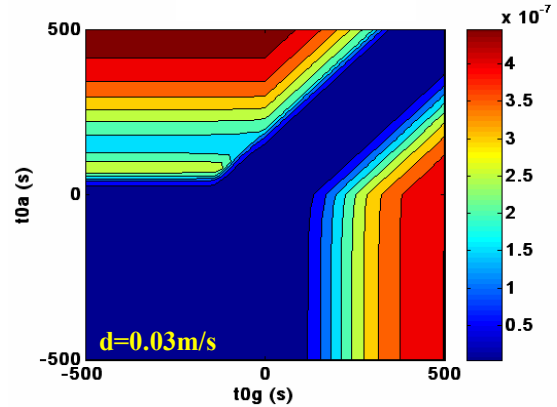


Figure 13. Worst-case $P_{LOI|fault}$ with d fixed.

Next, we use the above result to calculate $P(LOI|fault_k)$ with t_{0g} fixed at a value far behind t_{0a} , and vary t , t_{0a} and d . Figure 14 shows a 3-D plot of the resulting values of $P(LOI|fault_k)$. An interesting point to note from this plot is that as the fault magnitude (d) increases $P(LOI|fault_k)$ is worst when t_{0a} is small and time (t) approaches t_{0a} .

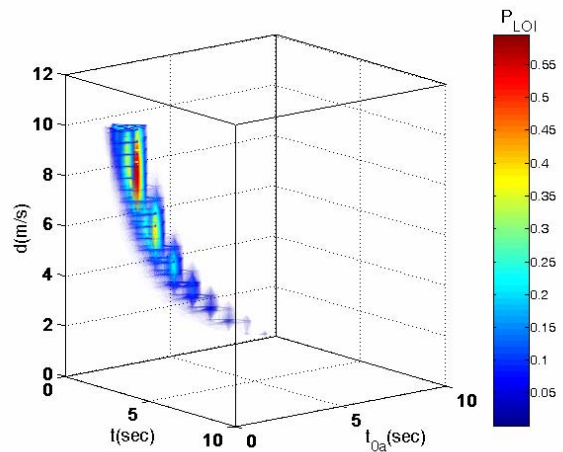


Figure 14. Worst case $P_{LOI|fault}$

To explain this result, consider again Figure 12 were the LGF filter starts well before the divergence onset (at $t = 0$) and the aircraft filter starts after the fault onset. Now, $P_{md|fault}$ improves as time increases. Hence, $P_{md|fault}$ is worst when the monitor response to fault has just started. i.e. for small t . On the other

hand, $P_{ev|fault}$ gets worse as $t \rightarrow t_{0a}$ and then improves for large values of t as the two filters approach the same steady state value. Therefore, the worst-case $P(LOI|fault_k)$ as shown in Figure 14 occurs at small values of t_{0a} with $t=t_{0a}$.

The total specified Time to Alert (TTA) for LAAS is 6 seconds, with 3 seconds allocated to the ground and 3 seconds for air. The preliminary TTA analysis in this paper does not yet address the issues related to transmission delays of signal if a fault is detected.

In order to carry out a preliminary TTA analysis for different fault magnitudes, we check the instances in which $P(LOI|fault_k) > 10^{-4}$ (the specified maximum probability of MI given the CCD fault) for each value of d , and the duration of such occurrences to verify if it is under 3 sec (allocation for ground). Figures 15 and 16 show the results.

Figure 15 shows the worst-case value (for any t , t_{0g} , and t_{0a}) of $P(LOI|fault)$ and the duration of time this probability exceeds 10^{-4} . It is clear from this figure, that there exist some occurrences of time in LOI of 3.5 sec, which exceeds the ground allocation to the time-to-alert (3 sec). The figure clearly shows that the result is on the edge of meeting the time-to alert requirement. This observation is further supported by the fact that the situation is easily remedied by reducing the monitor filter time constants from 30 sec to 29 sec. The result is shown in Figure 16.

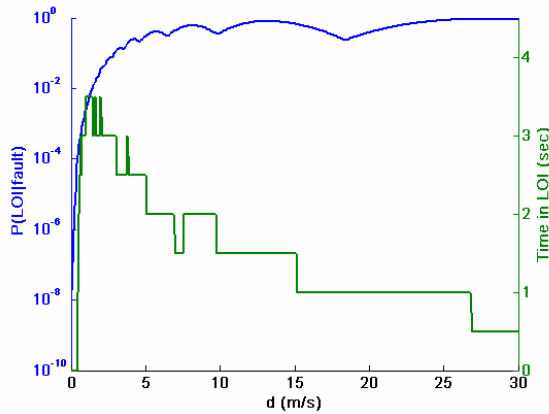


Figure 15. CCD LTI/LTV P(LOI) and Time in LOI

However, even this minor reduction in the filter time constant may not be necessary. The analysis is conservative because fault-free ionospheric divergence ranging errors are implicitly included in v_i in the development of equation (6). In reality, these errors will contribute directly to the divergence failure—i.e., slightly changing the effective value of d . The result is that the probability in equation (6) is conservatively computed in this analysis. A substantiation of these claims and a more detailed analysis of time-to-alarm problem will be the subjects of a future paper.

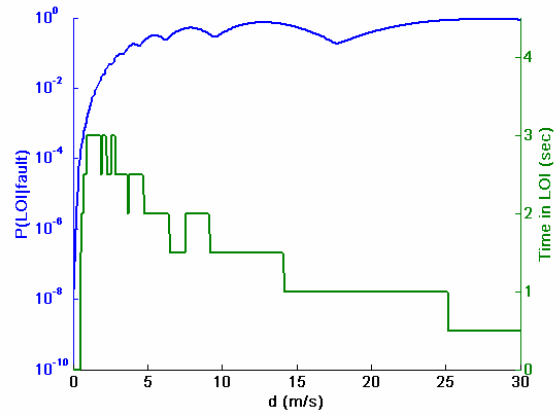


Figure 16. CCD LTI/LTV P(LOI) and Time in LOI

V. Conclusion

In this paper we have addressed the monitoring and integrity risk analysis for code minus carrier divergence GPS satellite faults. The paper suggests a 1st order LTV smoothing filter as a good choice for implementation at the aircraft and focuses the analysis on this implementation, but as other higher filter implementations are also possible the analysis approach was designed to be easily extendable to different filter implementations at the aircraft.

A CCD monitor to be implemented at the LAAS ground facility was designed. This monitor uses two 1st order LTI filters in series for divergence rate estimation, which is followed by a simple detection test. The paper gives the nominal values for monitor filter time constants. The detection threshold is set by computing the fault free test metric based on experimental data.

We also present a new direct approach to compute the probability of Loss of Integrity for a space segment failure and apply it to the CCD monitoring problem. Different aircraft and ground filter start times are explicitly accounted for. Preliminary analysis results show that LAAS integrity requirements are satisfied.

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